# Geochemistry of Melt Inclusions from the Fondo Riccio and Minopoli 1 Eruptions at Campi Flegrei (Italy)

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### MASTER OF SCIENCE

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#### ABSTRACT

Campi Flegrei is a large volcanic complex located west of the city of Naples, Italy. The area has been the site of volcanic activity for more than 60 ka and represents a potential volcanic hazard owing to the large local population. In this study, the geochemistry of the magma associated with two different eruptions at Campi Flegrei has been characterized, with the aim to identify geochemical trends that may help to predict the style and nature of future eruptions. Two eruptions of different age and eruptive style have been selected for study, Fondo Riccio (9.5 ka) and Minopoli 1 (11.1 ka). A scoria (CF-FR-C1) and a bomb (CF-FR-C2) were collected from the Fondo Riccio eruption, and two scoria samples were collected from Minopoli 1 (CF-Mi1-C1 and C2) eruption.

The pre-eruptive volatile content of magma plays an important role in the style of eruption and can be assessed from studies of melt inclusions (MI) contained in phenocrysts. Major and trace elements in Fondo Riccio MI show a wider variation compared to those in Minopoli 1 MI suggesting that the Fondo Riccio magma residence time was longer compared to the Minopoli 1 magma. Analyses of volatile contents in MI suggest that Fondo Riccio magma may have been more water-rich than Minopoli 1 magma, consistent with the more explosive character of this eruption compared to Minopoli 1. Trace element data suggest a combination of arc volcanic and upper continental crust magma as the source for the Fondo Riccio and Minopoli 1 eruptions.

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#### **CHAPTER 1: INTRODUCTION**

The geochemical evolution of an active volcanic system and identification of the parameters that play a role in determining the style of an eruption are of fundamental importance to understand the past behavior of a magmatic system and to forecast future behavior. Development of geochemical models for volcanic eruption forecasting require information on the volatile content of the magma before an eruption, because volatiles play a major role in controlling the nature and style of eruptive events (Anderson, 1976; Burnham, 1979). The exsolution and expansion of volatiles (especially H<sub>2</sub>O) provides the mechanical energy that drives explosive volcanic eruptions. The original volatile content of magma can be estimated by analyzing melt inclusions (MI) contained in phenocrysts (Anderson, 1974; Clocchiatti, 1975; Roedder, 1979; Belkin et al., 1985; Sobolev, 1990; Lowenstern, 1994; Anderson, 2003; Wallace, 2005). Moreover, MI may provide information concerning crystallization and mixing histories of magmas and also the conditions of primary melt generation and extraction (Roedder, 1984; Carroll and Holloway, 1994; Lowenstern, 1994; Sobolev, 1996; Marianelli et al., 1999; Danyushevsky et al., 2000; Frezzotti, 2001).

The Campi Flegrei volcanic complex is an active volcanic field in the Neapolitan area (Italy) that has experienced predominantly explosive volcanic activity for more than 60,000 years. The city of Pozzuoli lies close to the Solfatara crater (actually it was built on the deposits of numerous eruptions of the past 10 ka) while Naples, with 1.5 million inhabitants, is nearby, between Campi Flegrei and Vesuvius. The volcanic risk in this area is significant because of the large population and is a compelling reason to better

understand the evolution of the Campi Flegrei complex and the mechanisms that lead to explosive eruptions. We studied the products of two eruptions that occurred in Campi Flegrei during the First Cycle: Fondo Riccio (9.5-10.3 ka) and Minopoli 1 (10.3-11.5 ka). Fondo Riccio was an explosive strombolian eruption that occurred near the center of the Campi Flegrei caldera, whereas Minopoli was an explosive hydromagmatic eruption that occurred along the regional fault system in the northern portion of the Campi Flegrei caldera. Data from MI were used to constrain the evolution of major and volatile (H<sub>2</sub>O, Cl, S and F) element concentrations, with the goal to reconstruct the crystallization history of the magmas and assess pre-eruptive volatile abundances for the two eruptions.

#### **CHAPTER 2: GEOLOGICAL SETTING**

The Campi Flegrei Volcanic District lies in the Campanian Plain (CP), between the western side of the Southern Apennine Chain and the eastern border of the Tyrrhenian abyssal plain. Since late Miocene-early Pliocene, the Tyrrhenian Sea has been opening (Scandone, 1979; Doglioni, 1991) and the Calabrian arc has migrated to the SE following rollback of the subducted Ionian plate under Calabria (Selvaggi and Chiarabba, 1995; Piromallo and Morelli, 1997; Gvirtzman and Nur, 2000). Extension in the Tyrrhenian basin was accompanied by contemporaneous compression in the Apennine chain (Meletti et al., 2000). As a result of motions of the Tyrrhenian and Ionian blocks, the CP became a structural depression bordered by NW-SE and NE-SW trending faults (D'Argenio et al., 1973; Ippolito et al., 1975; Hippolyte et al., 1994). Geological, geophysical and petrologic evidence (Selvaggi and Amato, 1992; Serri et al., 1993; Peccerillo, 1999) suggest that subduction of oceanic lithosphere (from the relict Ionian basin) beneath the Apennines occurred concomitant with thinning of the continental lithosphere in the region of the Adriatic Sea, Sicily and North Africa.

The Campi Flegrei caldera (Fig, 1) is one of the most active volcanic systems in the Mediterranean region. The area is known for intense hydrothermal activity, frequent earthquakes and bradyseismic events that occurred between1969-1972 and 1982-1984. The area has been volcanically active for 60 ka (Pappalardo et al., 2002) and many studies have been devoted to understanding its activity (Di Girolamo et al., 1984; Rosi and Sbrana, 1987; Barberi et al., 1991; Pappalardo et al., 1999; De Vivo et al., 2001; Rolandi et al., 2003). At Campi Flegrei, numerous eruptions from multiple sources have produced lava and pyroclastic deposits, including several lava dome structures. Some authors (Rosi and Sbrana, 1987; Orsi et al., 1996) relate the origin of Campi Flegrei either to the eruption of the Campanian Ignimbrite (CI) (39 ka, De Vivo et al., 2001), or to the Neapolitan Yellow Tuff (NYT) (15 ka, Deino et al., 2004). An interpretation that considers the eruption of the CI to be a unique event originating in the Campi Flegrei caldera has been questioned by De Vivo et al. (2001) and Rolandi et al. (2003). These authors describe a sequence of eruptive events from fractures activated along the neotectonic Apennine fault system parallel to the Tyrrhenian coastline. These events, of ignimbritic origin, lasted from >300 ka to 19 ka and are not confined to a unique volcanic center in Campi Flegrei (Rosi and Sbrana, 1987; Orsi et al., 1996). Only the Neapolitan Yellow Tuff (NYT) (15ka, Deino et al., 2004) erupted within Campi Flegrei, whereas the CI (39 ka, De vivo et al., 2001) has a much wider source area (Rolandi et al., 2003). According to Pappalardo et al. (2002), the time between the CI and NYT eruptions is characterized by a large number of significantly less powerful events. Volcanism in this interval is poorly defined, primarily because of limited exposure due to cover by younger deposits, restriction to submarine exposure, and intense urbanization. Since the NYT eruption the edge of the caldera has been the site of at least 65 eruptions during three periods of activity (15.0–9.5; 8.6–8.2; and 4.8–3.8 ka). These eruptions were separated by quiescent periods marked by two widespread paleosoils (Di Vito et al., 1999). During each eruptive period, eruptions were separated by short-time intervals, on the order of tens of years. The last eruption in 1538 formed the Monte Nuovo cone (Di Vito et al., 1987) after 3.4 ka of quiescence.

Campi Flegrei eruptions were mostly explosive with variable degrees of magma/water interaction; only a few events were effusive. The volcanic products range from trachybasalt to alkali-trachyte and phono-trachyte, and are characterized by variable Sr-Nd-Pb-B isotope ratios (D'Antonio et al., 1999a; Pappalardo et al., 2002; Tonarini et al., 2004) and involve different magma types (i.e. from different sources and/or with different history) or source processes.

#### **CHAPTER 3: SAMPLES AND METHODS**

#### 3.1 Description of samples and petrography

Products from two eruptions that occurred in the First Epoch were studied. The Fondo Riccio eruption occurred at 9.5 ka from an eruptive center on the western side of the Gauro volcano, near the center of the Phlegraean caldera. The eruption was explosive, with strombolian character. The eruptive deposits are composed of coarse scoria beds with subordinate coarse-ash beds and overlay Montagna Spaccata tephra and are in turn overlain by paleosol A (Di Vito et al., 1999).

The Minopoli 1 eruption was a magmatic eruption, with hydromagmatic phases, that occurred at 11.1 ka, along a regional fault system in the northern sector of the Phlegraean caldera. The eruptive products are composed of alternating pumice lapilli fallout and mainly massive ash fallout, and less abundant cross laminated ash surge beds. The deposits overlay Pomici Principali and are overlain by Montagna Spaccata (Di Vito et al., 1999).

Sample CF-FR-C1 is a scoria and sample CF-FR-C2 is a lava bomb from the Fondo Riccio unit. Both samples are porphyritic latite, with abundant crystals in a glassy, vesicular groundmass, and were deposited between 9.5 and 10.3 ka (Di Vito et al., 1999). The bulk rock contains less than 10% phenocrysts, which include clinopyroxene, olivine and biotite with subordinate plagioclase and magnetite. In thin section clinopyroxene and plagioclase are found in small clots; these minerals also occur among microlites. Silicate MI in olivine and pyroxene phenocrysts from sample FR-C1 and in clinopyroxenes for

the sample FR-C2 were studied. For the Minopoli 1 unit, sample MI1-C1 is a scoria and sample MI1-C2 is a cohesive scoria sampled along the caldera rim, and was likely erupted during the final magmatic stage of the eruption. The deposits include scoria and ash layers. Both samples are porphyritic trachybasalt and range from 11.1 and 10.3 ka (Di Vito et al, 1999). The bulk rock contains about 20% phenocrysts, which include clinopyroxene, olivine and biotite. Silicate MI in olivine and pyroxene from Mi1-C1 and in clinopyroxene from MI1-C2 were studied.

The abundance of MI varies from crystal to crystal in the same sample. Melt inclusions consist of silicate glass, generally devitrified, with a shrinkage bubble and daughter crystals (usually apatite or Fe/Ti oxides). MI generally have elongated rectangular shapes and range from 30 to 80  $\mu$ m (most between 20 and 50  $\mu$ m) (Fig. 2). Rock samples were hand crushed and phenocrysts were hand picked, mounted on a glass slide, and doubly polished to improve the optical clarity during microscope heating experiments.

#### 3.2 Heating experiments

MI from Fondo Riccio and Minopoli 1 were partially or totally recrystallized when found and required heating and quenching to obtain a homogeneous glass. Heating experiments have been carried out at the Department of Geosciences at Virginia Tech (Blacksburg, USA) using the heating stage designed and built at the Vernadsky Institute, Moscow (Sobolev and Slutski, 1984). All the experiments were conducted in controlled (inert gas) atmosphere to avoid oxidation of the crystal according to the procedure described in Danyushevsky et al. (2000) and Lima (2000). The precision of temperature measurements is  $\pm$  3°C at 1200°C (measured by thermocouple), based on calibration using known melting points of Ag (938° C) and Au (1064°C). The inclusions were heated until all solid phases disappeared. Total homogenization (i.e. disappearance of the bubble) was not attempted to avoid decrepitation. The partial homogenization temperature (T<sub>h</sub>) thus represents a minimum estimate of the trapping temperature.

### 3.3 Electron Microprobe Analyses

After quenching, single crystals were mounted with epoxy on transparent polycarbonate rods and polished according to the technique described by Thomas and Bodnar (2002) to expose MI. MI were analyzed with a Cameca SX/50 Microprobe at University of Rome "La Sapienza" (IGAG-CNR, Rome, Italy) and at Virginia Tech (Blacksburg, USA). Analyses were performed at 15kV, using a current of 10nA with a defocused beam diameter of 10 µm and counting time 10 seconds, as recommended by Morgan and London (1996). Relative one-sigma precision is estimated to be 1 to 2 % for major elements and 5 to 10 % for minor elements. In each analytical run, alkalis were counted first, and no correction has been made for Na loss. Test runs made prior to the beginning of the analysis on synthetic and natural glass standards of known composition showed no significant alkali migration under the specified analytical conditions. Two points were analyzed in larger MI, whereas only one point was analyzed in smaller inclusions (<10µm). When no significant difference in composition was detected, an average of the two analyses was used. If the two analyses were significantly different the

results were discarded. If just one spot was possible, a comparison between the obtained data and data for other MI in the same sample was made to test for consistency. Note, however, that two or more analyses were obtained on ~90% of the MI. Host phenocrysts were analyzed at distances of about 20  $\mu$ m from MI with accelerating voltage of 15kV and beam current of 10 nA.

#### 3.4 Ion Microprobe Analyses

Selected MI were analyzed for H (reported as  $H_2O$ ), Li, Be, B, Rb, Sr, Y, Zr, Nb, Cs, Ce, Sm, Dm, Yb, Th and U by Secondary Ion mass Spectrometry (SIMS) at the Woods Hole Oceanographic Institution, using techniques detailed by Shimizu and Hart (1982) and Webster et al. (1996). Accelerating potential was 10kV and beam current was 1-2 nA. The inclusions were analyzed in one spot, five times each in depth profile mode. Precision and accuracy were monitored with NBS (National Bureau of Standards) reference glasses NBS 610. Results on the NBS glasses are similar and within 5% of the accepted values;  $H_2O$  concentrations are reproducible to  $\pm$  0.3 to 0.4 wt% and trace elements to 5 to 15% (for more details see Webster et al., 2001).

#### **CHAPTER 4: RESULTS AND DISCUSSIONS**

Fondo Riccio olivine compositions range from  $Fo_{84}$  to  $Fo_{87}$ , while Minopoli 1 olivines range from  $Fo_{77}$  to  $Fo_{78}$  (Fig. 3). Clinopyroxene compositions fall in the diopside-salite field (Wo<sub>44-48</sub>, Fs<sub>5-19</sub>) and Mg # (calculated based on total Fe) varies from about 73.4 to 88.6 for Fondo Riccio samples (with a compositional gap between 78 and 83) and from about 85.1 to 89.9 for Minopoli 1 samples. Clinopyroxenes are low-Ti, which is a characteristic of "HKS"- type lavas of the Roman Comagmatic Province (Cundari and Fergusson, 1982).

Homogenization temperatures of MI in clinopyroxene and olivine from the scoria of Fondo Riccio (FR-C1) average  $1135 \pm 3^{\circ}$ C and  $1155 \pm 3^{\circ}$ C, respectively, whereas MI in clinopyroxene from the bomb sample (FR-C2) average  $1159 \pm 3^{\circ}$ C. Homogenization temperatures of Minopoli 1 MI in clinopyroxenes average  $1132 \pm 3^{\circ}$ C and those in olivine average  $1145 \pm 3^{\circ}$ C.

The maximum temperature achieved during heating experiments does not equal the trapping temperature because a bubble still remained in the inclusions. MI compositions can be depleted in the host mineral components if the maximum temperature of heating is below the trapping temperature. Thus, the concentrations of most elements in the quenched MI may not reflect their original values in the trapped melt. However, ratios of elements that are incompatible in the host, and also concentrations of elements that are present at similar levels in the melt and the host, should not be significantly affected by over- or under-heating. Moreover, Fedele et al. (2003) have shown that compositions of MI obtained from inclusions that were heated until all of the solids had melted (with the

bubble still present) and quenched were in good agreement with melt compositions predicted by the MELTS program (see also Thomas et al, 2002).

Representative compositions of MI (average values) hosted in Fondo Riccio clinopyroxene and olivine are shown together with MI size, maximum temperature ( $T_{run}$ ) and host crystal Mg# in Table 1a (FR-C1) and 1b (FR-C2). Representative compositions of MI from Minopoli 1 are shown in Table 1c. All MI are characterized by analytical totals < 100. This is interpreted to reflect the presence of H<sub>2</sub>O in the MI, as confirmed by secondary ion mass spectrometry (SIMS) analyses that show average H<sub>2</sub>O contents between 3.3 - 6.9 wt% for Fondo Riccio samples and between 1.3-5.2 wt % for Minopoli 1 samples (Tables 1a-c).

MI from Fondo Riccio show a broad range for SiO<sub>2</sub> (46.8 - 59.3 wt%), CaO (3.66 - 12.4 wt%), Na<sub>2</sub>O (1.3-3.7 wt%) and K<sub>2</sub>O+Na<sub>2</sub>O (4.19 - 9.73 wt%). Conversely, MI from Minopoli 1 show narrow ranges for SiO<sub>2</sub> (49.6 -51.8 wt%), CaO (7.6 - 13.3 wt%), Na<sub>2</sub>O (1.0-1.8 wt%) and K<sub>2</sub>O+Na<sub>2</sub>O (4.19 - 9.73 wt%) as shown in Tables 1a-c.

On the total alkali-SiO<sub>2</sub> classification diagram (Fig.4; Le Bas et al., 1986) the Fondo Riccio bulk-rock composition is in the latite field. Fondo Riccio MI data show two trends: from latite to trachyte for MI in Fe-rich diopside (sample CF-FR-C1) and from trachybasalt to shoshonite for MI in both Mg-rich diopside (sample CF-FR-C2) and olivine (sample CF-FR-C1). The Minopoli 1 bulk-rock composition is in the trachyandesite field, whereas MI data show two different compositions: basalt for MI in clinopyroxenes from the lower unit (Mi1-C1) and trachybasalt for MI in olivine (Mi1-C1) and clinopyroxenes from the upper unit (Mi1-C2). Except for clinopyroxene in FR-C1, MI are generally less evolved than the corresponding host rock. In particular, MI in

olivine from CF-FR-C1 scoria show the same, less evolved composition as MI in clinopyroxene from CF-FR-C2 bomb sample. The Fondo Riccio bulk rock is depleted in K<sub>2</sub>O (not shown in the variation diagrams) and enriched in TiO<sub>2</sub> and Na<sub>2</sub>O compared with MI with similar Mg# hosted in clinopyroxene from the CF-FR-C1 scoria sample. Harker diagrams showing major element concentrations of MI as a function of SiO<sub>2</sub> show a wider range in composition in Fondo Riccio samples than Minopoli 1 (Fig. 5). MI in clinopyroxene and olivine from both samples show decreasing Al<sub>2</sub>O<sub>3</sub> (Fig. 5a) and Na<sub>2</sub>O (Fig. 5f) and increasing FeO<sup>tot</sup> (Fig. 5b), TiO<sub>2</sub> (Fig. 5c), MgO (Fig. 5d) and CaO (Fig. 5e) with decreasing SiO<sub>2</sub>. These trends are consistent with crystallization of olivine and clinopyroxene from alkali basalt magma. In particular, the trends described above suggest earlier crystallization of olivine in the CF-FR-C1 sample and clinopyroxene in the CF-FR-C2 sample compared with the timing of crystallization of clinopyroxene in the CF-FR-C1 scoria sample. The TiO<sub>2</sub> contents of these MI are consistent with the low Ticontents of lavas from the Roman Comagmatic Province. To explain Ti enrichment in the bulk rock compared with MI (whose compositions should represent the melt composition before the eruption), the studied rocks would have to contain about 60% groundmass and  $\approx$  40% phenocrysts (clinopyroxene, olivine, feldspar and biotite), which is a higher phenocryst abundant than that observed in thin sections. If we assume that the groundmass has a composition similar to MI in more evolved clinopyroxene, then TiO<sub>2</sub> bulk-rock contents are compatible with analyzed values (Table 1a-c).

It is more difficult to explain Na<sub>2</sub>O enrichment in the bulk rock, even though Na<sub>2</sub>O in the melt is expected to increase during crystallization (Na behaves as an incompatible element during magma fractionation). We also observe a greater Na

enrichment in Fondo Riccio bulk rock compared to Minopoli1 bulk rock. Hydrothermal activity is present in the vicinity of the Fondo Riccio eruption and, as explained for similar behavior at Vesuvius (Lima et al., in press), Na<sub>2</sub>O enrichment in the bulk rock most likely represents the effects of hydrothermal fluids. During explosive eruptions associated with decompression during magma ascent, the reaction of NaCl with H<sub>2</sub>O at low pressure (< 300 bars) becomes important and hydrolysis reactions produce HCl and NaOH that remain preferentially in the melt (Veksler, 2004). Note that the trend in Na enrichment observed here is the opposite of that observed during subsolidus hydrothermal alteration of melt inclusions in porphyry copper systems (Student and Bodnar, 2004).

The volatile content of MI as a function of SiO<sub>2</sub> for the two eruptions are shown in Fig.6. The SO<sub>2</sub> content of MI in clinopyroxene from scoria (C1) varies between 0.04 and 0.29 wt%; and from 0.25 to 0.40 wt% for MI in olivine from the scoria (C1) and from 0.13 to 0.34 wt% for MI in clinopyroxene from the bomb (C2). For Minopoli 1 MI, SO<sub>2</sub> varies from 0.09 to 0.11 wt% for clinopyroxene from C1, from 0.21 to 0.34 wt% for MI in olivine from C1, and from 0.31 to 0.37 wt% for MI in clinopyroxene from C2. The SO<sub>2</sub> concentration decreases with increasing SiO<sub>2</sub>, consistent with degassing during magma crystallization. MI hosted in clinopyroxene from FR-C1 shows little variation in S content as a function of SiO<sub>2</sub>. One Minopoli 1 MI containing approximately 48 wt.% SiO<sub>2</sub> and 0.91 wt% SO<sub>2</sub> falls well outside the range of all other values. This data point was not discarded because it represents the average of two nearly identical analyses on the same MI.

Decreasing  $P_2O_5$  with increasing SiO<sub>2</sub> (Fig. 6b) is consistent with apatite crystallization during magma evolution. No systematic variation in F is observable for the two samples (Fig. 6c).

Chlorine abundance (Fig. 6d) in the Fondo Riccio MI ranges from 0.32 to 0.72 wt % in MI in clinopyroxene from the scoria (mean of 0.54 wt%), whereas Cl shows much less variation in MI in clinopyroxene from the bomb sample (CF-FR-C2) (between 0.42 and 0.57 wt % with a mean of 0.49 wt%) and in MI in olivine from CF-FR-C1 sample (between 0.42 and 0.47 wt % with a mean of 0.44 wt %). Minopoli 1 MI show a smaller variation in Cl content compared to Fondo Riccio. Cl varies from 0.34 to 0.2 wt% in MI in clinopyroxenes from Mi1-C1 (mean 0.27 wt%), 0.43±0.01 wt% for MI in olivine from Mi1-C1 and from 0.39 to 0.42 wt.% for MI in clinopyroxene from Mi1-C2. These results suggest that the magma that generated the Fondo Riccio scoria phenocrysts was not affected by Cl loss. The Cl content is slightly higher in more evolved MI than in less evolved MI for Fondo Riccio samples (Fig. 6d), whereas such a trend is not observed for Minopoli 1 MI.

The average H<sub>2</sub>O content, measured by EMPA, is consistent with the H<sub>2</sub>O obtained by SIMS (Fig.7). In general, less evolved MI have higher concentrations of water and Fondo Riccio MI are enriched in H<sub>2</sub>O compared to Minopoli 1 MI (even though the concentrations overlap). This may explain the more explosive character of the Fondo Riccio eruption compared to the Minopoli 1 eruption.

In order to better understand the origin of the magma that fed these two eruptions, trace element systematics have been examined. Compositions of MI from both eruptions have B and Be concentrations that fall close to the B/Be = 4 line that divides rocks that

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originate in volcanic arc environments and those that are sourced in an ocean island volcanic setting (Fig. 8). Rb versus Y+Nb systematics of MI straddle the boundary between rocks associated with "within plate" volcanism and those generated in volcanic arc settings (Fig. 9). Y/Nb versus Zr/Nb systematics of MI are consistent with magmas generated in the upper continental crust, but also point towards the island arc environment (Fig. 10). An upper continental crust origin for the MI is also suggested by the Sr, Rb, Th, Nb, Ce, Zr, Sm, Y and Yb spider diagram (Pearce, 1984) (Fig.11).

The relationship between Rb and Cs for MI from Somma-Vesuvius (SV) and the Campanian Ignimbrite (CI) (shaded areas) from Webster et al., 2003 are compared with MI from this study in Fig. 12. The data suggest that Fondo Riccio and Minopoli 1 eruptive products were more likely generated from a magma similar to that which fed Somma Vesuvius rather than that which produced the Campanian Ignimbrite rocks.

The trace element data presented above are consistent with the tectonic setting of the Neapolitan area, and with the possible involvement of slab-derived fluids, as suggested by Piochi et al. (2005).

#### **CHAPTER 5: CONCLUSIONS**

MI from Fondo Riccio and Minopoli 1 show systematic variations in composition compared to the bulk-rock compositions. Major and trace element composition of MI are consistent with an evolving magma chamber in which olivine and clinopyroxenes are crystallizing. The composition of MI in olivine from Fondo Riccio scoria (CF-FR-C1) is similar to MI in clinopyroxene from the bomb (CF-FR-C2). Except for clinopyroxene in Fondo Riccio C1, MI are generally less evolved than the corresponding host rock. Major and trace elements in Fondo Riccio MI show a wider variation compared to those in Minopoli 1 MI. This could be interpreted to indicate that the Fondo Riccio magma residence time was longer compared to the Minopoli 1 magma. Na<sub>2</sub>O enrichment in the Fondo Riccio bulk rock represents the effects of hydrothermal activity in the volcanic system before and during the explosive eruptions (see, De Vivo and Lima, in press).

The concentration of SO<sub>2</sub> and H<sub>2</sub>O is higher in less evolved MI, whereas Cl is higher in more evolved MI. F shows no obvious variation as a function of SiO<sub>2</sub>. The generally higher volatile contents of Fondo Riccio MI are consistent with the more explosive character of this eruption compared to Minopoli 1. Trace element data suggest a combination of arc volcanic and upper continental crust magma as the source for the Fondo Riccio and Minopoli 1 eruptions. Major and trace element systematics suggest that these two eruptions, which occurred at different times, were sourced from a single batch of magma that has evolved with time. The smaller range in compositions of the Minopoli 1 MI compared to the Fondo Riccio samples suggest a shorter residence time for Minopoli compared to Fondo Riccio magma.

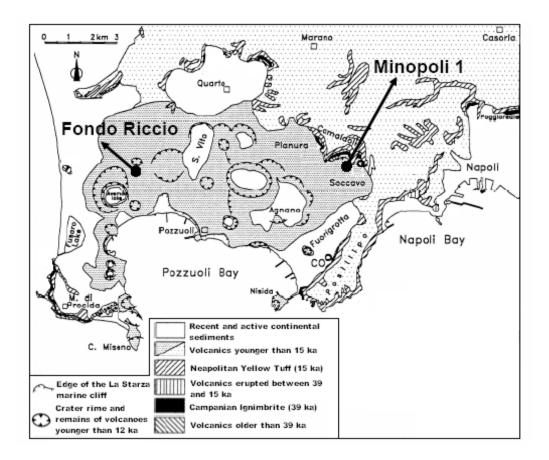


Fig.1

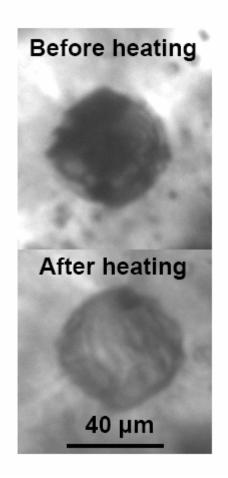


Fig. 2

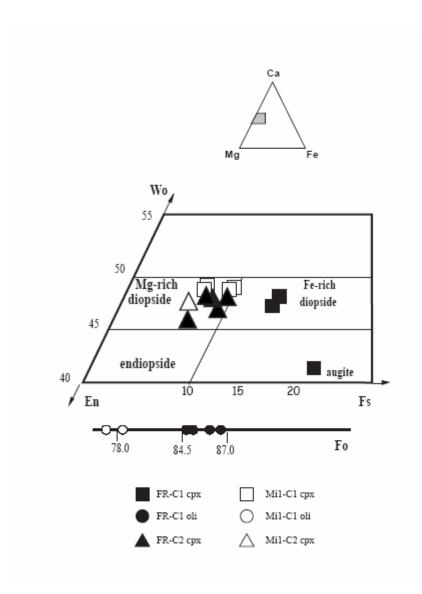


Fig.3

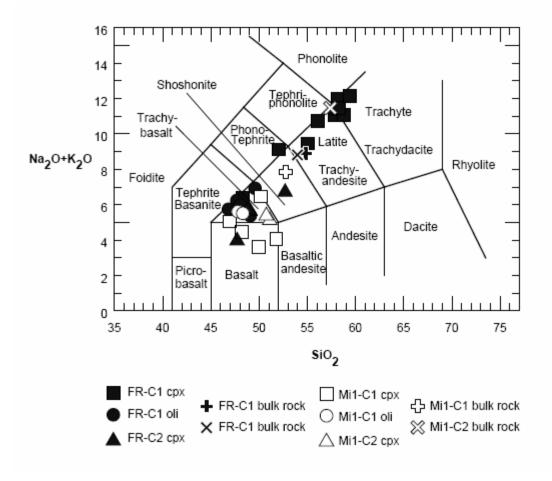


Fig 4

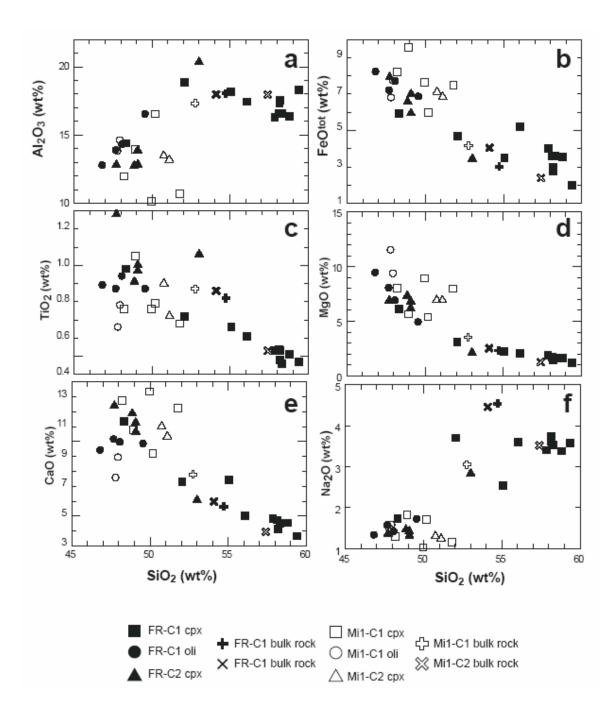


Fig.5

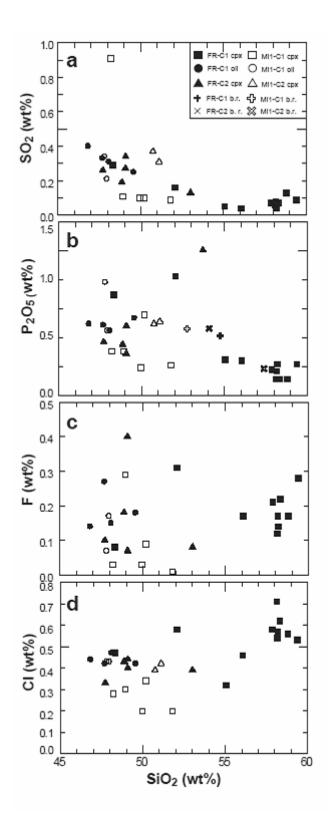


Fig.6

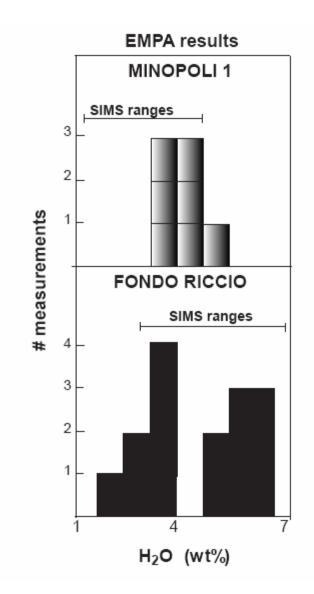


Fig.7

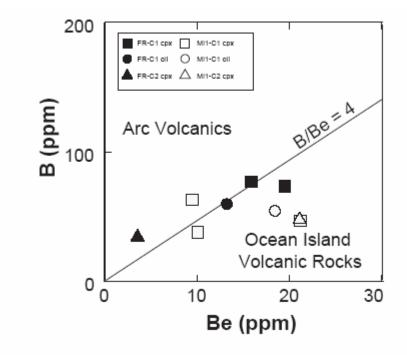


Fig.8

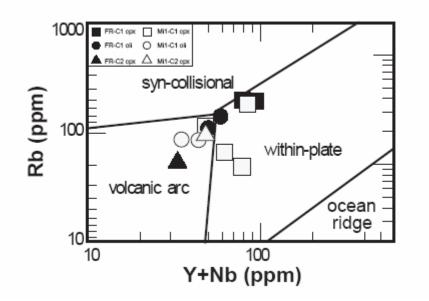


Fig.9

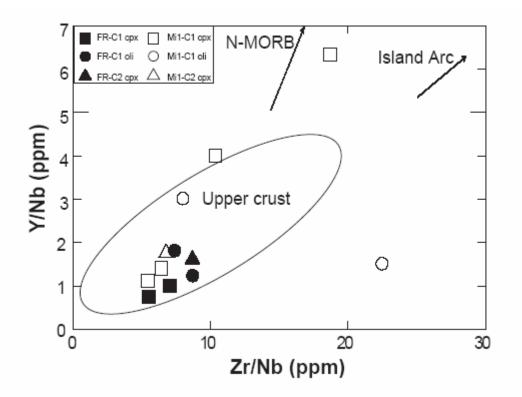


Fig.10

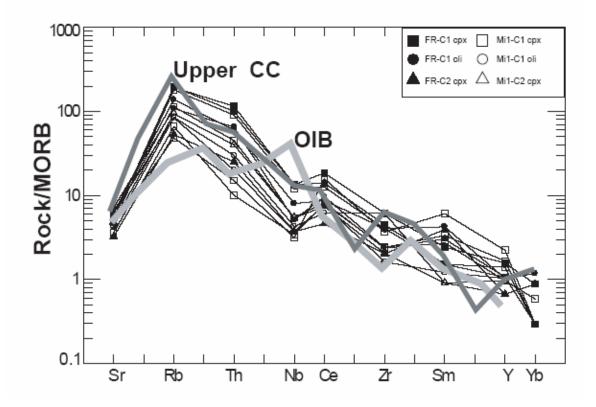


Fig.11

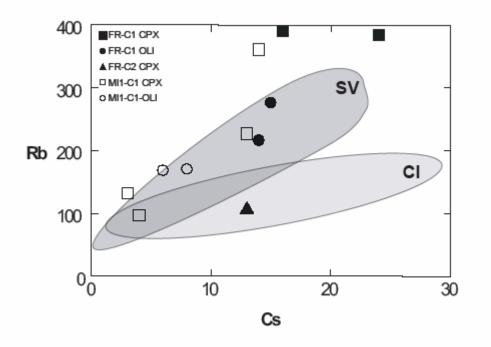


Fig.12

#### **FIGURE CAPTIONS**

- **Figure 1.** Schematic map of the Campi Flegrei area showing sample locations (modified after Pappalardo et al., 1999).
- Figure 2. Change in appearance of two crystallized silicate melt inclusions hosted in clinopyroxene in Fondo Riccio scoria as a result of heating from room temperature (22°C) to the homogenization temperature (1137°C). After heating to 1137°C the inclusion contains only a vapor bubble. The MI is approximately xx μm in diameter.
- Figure 3. Classification diagrams for pyroxene and olivine phenocrysts (Morimoto, 1988). Filled squares = compositions of MI hosted in clinopyroxene from FR-C1 (scoria); filled circles = composition of MI hosted in olivine from FR-C1 (scoria); filled triangles = compositions of MI hosted in clinopyroxene from FR-C2 (bomb); open squares = compositions of MI hosted in clinopyroxene from Mi1-C1; open circles = compositions of MI hosted in olivine from Mi1-C1; open triangles = composition of MI hosted in clinopyroxenes from Mi1-C1; open triangles = composition of MI hosted in clinopyroxenes from Mi1-C2.
- Figure 4. Total alkali-silica diagram (Le Bas et al., 1986) showing compositions of MI hosted in clinopyroxene and olivine from Fondo Riccio and Minopoli 1.Bulk rock compositions are also shown.
- Figure 5. Harker variation diagrams for MI plotted versus SiO<sub>2</sub>: (a). Al<sub>2</sub>O<sub>3</sub>; (b). CaO;
  (c) Na<sub>2</sub>O.; (d). P<sub>2</sub>O<sub>5</sub>; (e) TiO<sub>2</sub>; (f) FeO<sup>tot</sup>.

- Figure 6. Water contents of MI obtained by EMPA and SIMS.
- **Figure 7**. Volatile concentrations in MI from Fondo Riccio and Minopoli 1 plotted versus SiO<sub>2</sub> concentration.
- **Figure 8.** Relationship between B and Be for Fondo Riccio and Minopoli 1 MI. The diagonal line corresponds to B/Be = 4 and divides arc volcanic magmas from ocean island volcanic magmas. Both samples fall along the boundary line, suggesting a mixed arc volcanic and ocean island origin. Two data points for samples Mi1-C1-p6 M2 and FR-C1-o4 M1 have not been plotted.
- Figure 9. Relationship between Rb and Y+Nb for Fondo Riccio and Minopoli 1 MI.MI compositions are near the boundary between the volcanic arc and "within plate" fields, suggesting a mixed magma source (see also Fig.1 in Piochi et al., 2005).
- Figure 10. Relationship between Y/Nb and Zr/Nb for Fondo Riccio and Minopoli 1 MI. Compositions of MI from both samples suggest an upper crustal magma source.
- Figure 11 Spider diagram for Fondo Riccio and Minopoli 1 MI (Pearce, 1983). The heavier lines show the upper continental crust (Upper CC) and the oceanic island basalt (OIB) trends.
- Figure 12 Relationship between Rb and Cs for Fondo Riccio and Minopoli 1 MI. The data suggest that the feeding mechanism of these two eruptions is more similar to the Somma-Vesuvius (SV) magma source than the Campanian Ignimbrite source.

### Table 1a

Sample	FR-C1	FR-C1	FR-C1	FR-C1	FR-C1	FR-C1	FR-C1	FR-C1	FR-C1	FR-C1	FR-C1	FR-C1	FR-C1	FR-C1	FR-C1
MI label	p6 M1	p6 M2	p6 M4	p6 M5	p6 M6	p6 M7	p6 M8	p6 M9	p11 M1	p12 M1	p12 M2	o1 M1	o2 M1	o4 M1	o6 M1
Size (µm)	75	80	35	30	60	55	75	55	85	50	130	70	50	60	50
Host xI	срх	срх	срх	срх	срх	срх	срх	срх	срх	срх	срх	ol	ol	ol	ol
T (quen)°C		1126	1126	1126	1126	1126	1126	1126	1145	1139	1139	1172	1150	1175	1135
Host Mg#	74.41	73.37	73.68	75.16	76.11	83.17	78.04	86.12	83.17	78.04	86.12	86.77	86.56	84.95	84.70
Host xI con	npositi	ion													
SiO <sub>2</sub>	49.93	50.24	50.24	50.36	50.36	50.36	50.74	50.41	51.68	49.01	49.80	40.15	40.02	40.74	39.77
TiO <sub>2</sub>	0.62	0.63	0.63	0.63	0.63	0.63	0.59	0.53	0.32	0.77	0.67	0.05	0.00	0.04	0.00
Al <sub>2</sub> O <sub>3</sub>	3.21	3.35	3.35	3.02	3.02	3.02	3.27	3.17	3.02	4.13	3.63	0.01	0.01	0.03	0.02
FeO	8.52	8.71	8.71	8.71	8.71	8.71	8.17	7.99	5.62	7.43	6.49	13.06	12.90	13.89	14.71
MnO	0.44	0.48	0.48	0.40	0.40	0.40	0.39	0.39	0.11	0.24	0.21	0.23	0.25	0.28	0.30
MgO	13.90	13.47	13.47	13.68	13.68	13.68	13.87	14.27	15.59	14.83	15.23	48.06	46.62	44.00	45.71
CaO	22.45	22.41	22.41	22.25	22.25	22.25	22.04	21.92	22.78	21.91	22.37	0.36	0.32	0.30	0.31
Na <sub>2</sub> O	0.36	0.41	0.41	0.37	0.37	0.37	0.34	0.37	0.17	0.28	0.22	0.03	0.01	0.01	0.01
K <sub>2</sub> O	0.02	0.00	0.00	0.02	0.02	0.02	0.01	0.00	0.01	0.01	0.02	0.02	0.01	0.00	0.03
P <sub>2</sub> O <sub>6</sub>	0.02	0.02	0.02	0.02	0.02	0.02	0.03	0.00	0.01	0.01	0.01	0.02	0.01	0.00	0.02
Total	99.57	99.99	99.99	99.67	99.67	99.67	99.50	99.09	99.36	98.71	98.70	102.14	98.82	99.38	100.90
Wo	07.0	07.0	07.0	07.0	07.0	07.0	07.0	07.0	E 4 7	50.5	E0 5				
	27.3	27.3	27.3 56.5	27.3 56.5	27.3	27.3	27.3	27.3	51.7	50.5	50.5 34.4	0.6	0.5	0.5 75.3	0.5
En Fs	56.5 16.2	56.5 16.2	50.5 16.2	50.5 16.2	56.5 16.2	56.5 16.2	56.5 16.2	56.5 16.2	35.4 13.0	34.4 15.1	34.4 15.1	77.9 21.5	77.6 21.9	75.3 24.2	74.9 24.6
				10.2	10.2	10.2	10.2	10.2	15.0	19.1	19.1	21.0	21.0	27.2	24.0
	Melt inclusions composition Electron Microprobe														
	-		50.40	50.00	50.48	50.00	50.00	E 0 0 0	50.00	EE 0.0	40.00	47.80	48.00	40.00	40.57
SiO <sub>2</sub>		58.12				58.30			52.06	55.06	48.33	47.69		48.09	49.57
TiO <sub>2</sub>	0.53	0.54	0.48	0.51	0.53	0.46	0.47	0.61	0.72	0.66	0.98	0.87	0.89	0.94	0.87
Al <sub>2</sub> O <sub>3</sub>	16.32	16.61	17.55	16.35	17.33		18.32		18.88	18.19	14.39	13.88	12.77	14.31	16.52
FeO <sup>tot</sup>	4.01	3.58	2.77	3.54	2.97	3.61	1.98	5.21	4.68	3.51	5.94	7.21	8.22	7.71	6.86
MnO	0.26	0.23	0.16	0.32	0.14	0.16	0.13	0.07	0.15	0.06	0.11	0.12	0.18	0.17	0.18
MgO	1.91	1.74	1.44	1.64	1.47	1.66	1.21	2.05	3.08	2.25	6.11	8.05	9.45	6.92	4.91
CaO	4.81	4.74	4.14	4.54	4.16	4.51	3.66	5.02	7.32	7.42	11.35	10.15	9.42	9.96	9.86
Na <sub>2</sub> O	3.40	3.61	3.60	3.38	3.73	3.53	3.57	3.60	3.70	2.54	1.73	1.57	1.33	1.41	1.72
K₂O	7.65	7.78	8.18	7.69	8.25	7.94	8.57	7.12	5.42	6.92	4.64	4.67	4.41	4.58	5.21
P <sub>2</sub> O <sub>6</sub>	0.22	0.14	0.27	0.14	0.21	0.14	0.27	0.30	1.03	0.31	0.87	0.61	0.62	0.56	0.67
NiO	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
SO <sub>2</sub>	0.07	0.06	0.04	0.13	0.08	0.07	0.09	0.04	0.16	0.05	0.29	0.33	0.40	0.31	0.25
F	0.21	0.12	0.14	0.17	0.17	0.22	0.28	0.17	0.31	0.00	0.08	0.27	0.14	0.15	0.18
CI Total	0.58 97.83	0.71 97.96	0.57 97.52	0.56 97.76	0.54	0.62 97.76	0.53 98.46	0.46	0.58 98.09	0.32 97.27	0.47 95.29	0.42 95.85	0.44	0.47 95.58	0.42 97.22
1		87.80	87.02	81.10	81.14	87.70	80.40	80.10	80.08	81.21	80.28	80.60	80.06	80.08	81.22
lon Microp H₂O (wt%)		8.09					2.00					4.40		2.64	
	-	6.96 28	-	-	-	-	3.66 26	-	-	-	-	4.13 15	-	3.61 11	-
Li ppm Be	-	28 13	-	-	2	-	20 16	-	-	-	-	15	-	52	-
B		62	-	-	-	-	59	-	-	-	-	39	-	55	-
Rb	-	316	-	-		-	316	-	-	-	_	218	-	276	-
Sr		627	-	-	-	-	614	-	-	-	-	1043	-	902	-
Y	-	53	-	-	-	-	35	-	-	-	-	45	-	43	-
Zr	-	364	-	-	-	-	262	-	-	-	-	176	-	292	-
Nb	-	49	-	-	-	-	48	-	-	-	-	23	-	35	-
Cs	-	25	-	-	-	-	16	-	-	-	-	19	-	19	-
Ce	-	172	-	-	-	-	128	-	-	-	-	112	-	98	-
Nd	-	47	-	-	-	-	32	-	-	-	-	38	-	97	-
Sm	-	1	-	-	-	-	1	-	-	-	-	2	-	3	-
Dy	-	4	-	-	-	-	<1	-	-	-	-	2	-	3	-
Yb Th	-	1	-	-	-	-	<1 25	-	-	-	-	2 21	-	2	-
U	-	29 12	-	-	-	-	25 7	-	-	-	-	21	-	19 9	-
5	-	14	-	-	-	-	1	-	-	-	-	0	-	ð	-

# Table 1b

Sample	FR-C2	FR-C2	FR-C2	FR-C2	FR-C2
MI label	p2 M1	p4 M1	p5 M1	p6 M1	p7 M1
Size (µm)	20	40	80	100	40
Host xI	срх	срх	срх	срх	срх
T (quen)°C	1176	1135	1165	1152	1164
Host Mg#	88.68	86.94	88.57	84.98	84.62
Host xI comp	osition				
SiO <sub>2</sub>	53.70	52.03	52.11	51.25	51.67
TiO <sub>2</sub>	0.36	0.44	0.39	0.60	0.54
Al <sub>2</sub> O <sub>3</sub>	2.09	2.48	2.11	2.79	3.21
FeO	3.76	4.38	3.93	5.06	5.00
MnO	0.10	0.14	0.08	0.17	0.12
MgO	16.53	16.36	17.09	16.07	15.44
CaO	23.25	23.01	23.00	22.47	23.02
Na <sub>2</sub> O	0.11	0.11	0.09	0.15	0.12
K <sub>2</sub> O	0.00	0.01	0.02	0.01	0.00
P <sub>2</sub> O <sub>6</sub>	0.02	0.04	0.01	0.06	0.01
Total	100.07	99.05	98.94	98.67	99.20
14/-	50.0	FC 4	50.0	54.0	50.0
Wo En	53.3 37.9	52.4 37.3	52.2 38.8	51.3 36.7	52.8 35.4
Fs	37.9		38.8 9.1		
		10.3		11.9	11.7
Melt inclusio			1		
Electron Mic					
SiO <sub>2</sub>	49.09	49.08	48.87	47.72	53.01
TiO <sub>2</sub>	1.00	0.97	0.91	1.28	1.06
Al <sub>2</sub> O <sub>3</sub>	12.84	13.86	12.76	12.81	20.34
FeO <sup>tot</sup>	7.00	5.94	6.57	7.92	3.44
MnO	0.10	0.15	0.15	0.28	0.16
MgO	6.81	6.17	7.33	6.90	2.12
CaO	10.65	11.27	11.88	12.40	6.06
Na <sub>2</sub> O	1.31	1.42	1.46	1.36	2.83
K₂O	4.07	4.20	4.23	2.60	7.83
P <sub>2</sub> O <sub>6</sub>	0.60	0.36	0.44	0.46	1.36
NiO	0.00	0.00	0.00	0.00	0.00
SO <sub>2</sub>	0.34	0.27	0.19	0.26	0.13
F	0.07	0.40	0.18	0.10	0.08
CI	0.44	0.40	0.43	0.33	0.39
Total	94.31	94.49	95.40	94.42	98.81
lon Micropro					
H <sub>2</sub> O	3.25	-	-	-	-
Li	11	-	-	-	-
Be	3	-	-	-	-
B	24	-	-	-	-
Rb Sr	98 431	-	-	-	-
Sr Y	25	-	-	-	-
Zr	136	-	-	-	-
Nb	16	-	-	-	-
Cs	15	-	-	-	-
Ce	48	-	-	-	-
Nd	16	-	-	-	-
Sm	3	-	-	-	-
Dy	1	-	-	-	-
Yb	1	-	-	-	-
Th	7	-	-	-	-
U	6	-	-	-	-

# Table 1c

Sample	Mi 1-C1	Mi 1-C1	Mi 1-C1	Mi 1-C1	Mi 1-C1	Mi 1-C1	Mi 1-C1	Mi 1-C2	Mi 1-C2		
MI label	p2 M1	p3 M1	p6 M1	p6 M2	p8 M1	o5 M1	o6 M1	p1 M1	p1 M2		
Size (µm)	50	60	50	70	40	60	55	45	50		
Host xl	срх	cpx	срх	срх	cpx	olivine	olivine	cpx	срх		
T (quen)°C	1122	1149	1156	1156	1108	1120	1145	1122	1122		
Host Mg#	86.64	88.59	85.10	86.10	86.29	86.68	85.68	89.89	89.89		
	Host xI composition										
SiO <sub>2</sub>	52.39	52.96	52.52	52.52	52.78	39.38	38.96	52.67	52.67		
TiO <sub>2</sub>	0.41	0.33	0.43	0.43	0.36	0.00	0.01	0.25	0.25		
Al <sub>2</sub> O <sub>3</sub>	3.03	2.19	3.35	3.35	2.78	0.02	0.02	1.68	1.68		
FeO	4.36	3.72	4.97	4.97	4.58	12.66	13.70	3.37	3.37		
MnO	0.14	0.10	0.10	0.10	0.11	0.20	0.22	0.09	0.09		
MgO	15.86	16.20	15.92	15.92	16.18	46.22	45.99	16.82	16.82		
CaO	22.75	23.39	21.52	21.52	22.53	0.28	0.27	23.53	23.53		
Na <sub>2</sub> O	0.14	0.13	0.15	0.15	0.13	0.02	0.04	0.15	0.15		
K₂O	0.01	0.03	0.02	0.02	0.01	0.03	0.01	0.01	0.01		
P <sub>2</sub> O <sub>6</sub>	0.02	0.04	0.00	0.00	0.00	0.00	0.02	0.00	0.00		
Total	99.11	99.07	98.99	98.99	99.45	98.82	99.22	98.56	98.56		
Wo	52.8	53.9	50.6	50.6	51.9	0.5	0.4	53.7	53.7		
En	36.8	37.3	37.5	37.5	37.3	77.9	76.4	38.4	38.4		
Fs	10.4	8.8	11.9	11.9	10.8	21.7	23.1	7.9	7.9		
Melt inclusio	ns comp	osition									
Electron Micr	roprobe										
SiO <sub>2</sub>	46.94	48.20	51.79	49.96	50.19	47.81	47.96	51.11	50.75		
TiO <sub>2</sub>	1.05	0.76	0.68	0.76	0.79	0.66	0.78	0.72	0.90		
Al <sub>2</sub> O <sub>3</sub>	13.97	11.98	10.70	10.15	16.54	13.81	14.61	13.18	13.47		
FeO <sup>tot</sup>	9.58	8.20	7.50	7.66	5.99	6.80	7.73	6.84	7.11		
MnO	0.29	0.10	0.21	0.15	0.17	0.12	0.08	0.10	0.17		
MgO	5.65	8.00	7.98	8.93	5.40	11.53	9.37	6.92	6.93		
CaO	10.80	12.75	12.23	13.35	9.20	7.57	8.94	10.34	11.01		
Na <sub>2</sub> O	1.82	1.30	1.16	1.03	1.71	1.50	1.57	1.24	1.30		
K <sub>2</sub> O	3.24	3.17	2.91	2.58	4.76	4.11	3.92	3.88	4.07		
P <sub>2</sub> O <sub>6</sub>	0.38	0.38	0.26	0.24	0.70	0.98	0.56	0.64	0.62		
NiO	0.14	0.05	0.04	0.00	0.02	0.08	0.05	0.00	0.00		
SO <sub>2</sub>	0.11	0.91	0.09	0.10	0.10	0.34	0.21	0.31	0.37		
F	0.29	0.03	0.01	0.03	0.09	0.07	0.17	0.00	0.00		
CI	0.30	0.28	0.20	0.20	0.34	0.43	0.43	0.42	0.39		
Total	94.56	96.11	95.74	95.14	96.00	95.81	96.38	95.68	97.06		
Ion Micropro											
H₂O	5.15	-	1.93	1.29	5.28	3.89	4.40	5.15	-		
Li	21	-	11	12	24	7	18	21	-		
Be B	14	-	7 27	77 320	7 44	12 36	1	14 47	-		
B Rb	31 97	-	27 324	320 122	44 215	36	32 167		-		
Sr	97 667	-	324 614	122 480	215 765	168 607	167 889	176 646	-		
ər Y	92	-	52	63	37	27	42	37	-		
Zr	266	-	262	160	162	437	116	139	-		
Nb	14	-	48	14	23	19	16	22	-		
Cs	6	-	16	4	16	10	8	93	-		
Ce	180	-	128	83	88	69	87	79	-		
Nd	180	-	32	74	63	61	75	53	-		
Sm	4	-	1	2	2	1	1	1	-		
Dy	2	-	<1	1	1	<1	<1	<1	-		
Yb	<1	-	1	<1	1	<1	<1	<1	-		
Th U	5 1	-	25 7	3 2	14 4	9 3	6 <1	13 <1	-		
0	1	-	1	2	4	3	2	51	-		

Table 2

		CF-FR-C1	RE CF-FR-C1	CF-FR-C2	CF-MI1-C1	CF-MI1-C2
SiO <sub>2</sub>	wt%	54.76	54.68	54.10	52.76	57.40
Al <sub>2</sub> O <sub>3</sub>	wt%	17.99	18.11	17.98	17.32	17.97
Fe <sub>2</sub> O <sub>3</sub>	wt%	7.26	7.23	7.61	7.44	4.87
MgO	wt%	2.33	2.31	2.53	3.53	1.26
CaO	wt%	5.62	5.64	5.97	7.79	3.94
Na <sub>2</sub> O	wt%	4.49	4.59	4.46	3.05	3.52
K <sub>2</sub> O	wt%	4.35	4.34	4.35	4.79	7.94
TiO <sub>2</sub>	wt%	0.82	0.82	0.86	0.87	0.53
P <sub>2</sub> O <sub>5</sub>	wt%	0.52	0.51	0.58	0.58	0.23
MnO	wt%	0.13	0.13	0.13	0.12	0.12
Cr <sub>2</sub> O <sub>3</sub>	wt%	0.03	0.03	0.01	0.01	0.02
Ni	ppm	12.00	13.00	6.00	19.00	6.00
Sc	ppm	10.00	10.00	11.00	15.00	5.00
LOI	wt%	1.20	1.10	0.90	1.30	1.80
TOT/C TOT/S	wt% wt%	0.01	0.01	0.01	0.05	0.03
SUM	wt%	99.50	99.49	99.48	99.56	99.59
FeO*	wt%	3.2	3.2	4.05	4.17	2.39
Mo	ppm	2.40	2.5	2	1.6	0.5
Cu	ppm	7.40	6.6	7.5	15.3	3.8
Pb	ppm	6.50	6.3	7.2	10.8	3.9
Zn	ppm	62.00	60	73	44	9
Ni	ppm	5.30	5	3	7.6	2.9
As	ppm	4.70	4.8	5.7	3.8	1.9
Cd	ppm	<.1	<.1	<.1	0.1	<.1
Sb	ppm	0.20	0.2	0.2	0.1	0.1
Bi	ppm	0.10	0.1	<.1	0.1	0.1
Ag	ppm	<.1	<.1	<.1	<.1	<.1
Au	ppm	1.00	<.5	2.4	1.6	2.5
Hg	ppm	<.01	<.01	0.01	0.01	<.01
TI	ppm	0.50	0.5	0.2	0.5	0.1
Se Ba	ppm	<.5	<.5 1654	<.5	<.5	<.5
Ве	ppm	1719.20 7.00	9	1712.1 7	1715.2 7	1048.9 10
Со	ppm	15.90	16.4	17.2	21.5	7.6
Cs	ppm	16.90	16.7	16.2	11.8	16.1
Ga	ppm ppm	20.10	19.6	20.7	19.4	18.1
Hf	ppm	6.40	6.5	6.4	5	6.8
Nb	ppm	35.00	33.5	33.2	26.7	42.1
Rb	ppm	274.80	269.6	309.6	238.5	305.1
Sn	ppm	3.00	208.0	3	3	4
Sr	ppm	903.00	889.4	919.8	946	724.7
Та	ppm	1.80	1.8	1.8	1.4	2.3
Th	ppm	22.10	24.2	22.8	18.1	29.7
U	ppm	6.80	6.8	6.1	5.2	9
v	ppm	193.00	190	209	208	110
w	ppm	4.00	4	3.5	4.6	6.5
Zr	ppm	233.20	229.6	229.2	182	267.9
Y	ppm	32.50	31.8	33.4	28.7	32.1
La	ppm	64.10	63.6	63.7	56	74.3
Ce	ppm	123.60	121.7	123.8	109.7	139.6
Pr	ppm	13.40	13.39	13.74	12.37	14.82
Nd	ppm	53.90	53.3	56.8	49.2	57
Sm	ppm	10.10	10	10.5	9.7	9.9
Eu	ppm	2.37	2.28	2.5	2.24	2.27
Gd	ppm	7.66	7.37	7.78	6.99	7.14
ТЬ	ppm	1.08	1.05	1.1	1.03	1.04
Dy	ppm	5.89	5.85	5.97	5.47	5.65
Ho	ppm	1.08	1.06	1.07	0.99	1.07
Er	ppm	2.89	2.9	2.95	2.68	2.95
Tm	ppm	0.45	0.42	0.44	0.38	0.46
Yb	ppm	2.48	2.6	2.63	2.41	2.65
Lu Lu	ppm v dichro	0.41 mate tritation	0.41	0.42	0.36	0.41
- reo b	y ucrii 0	mate untation				

### **TABLE CAPTIONS**

- Table 1 MI composition determined by EMPA and SIMS. \*T(run) = maximum temperature of heating experiment. Cpx = clinopyroxene host crystal. Ol = olivine host crystal: (a). Fondo Riccio C1; (b). Fondo Riccio C2; (c). Minopoli 1 (C1-C2).
- **Table 2**Fondo Riccio and Minopoli 1 bulk rock compositions determined by XRF.

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